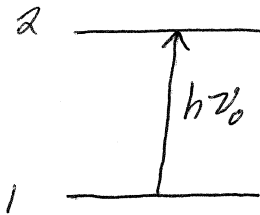
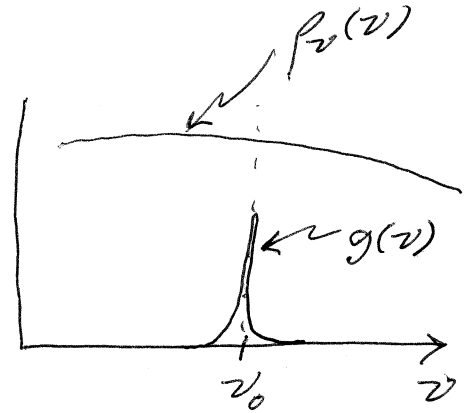


Transition Rates for Monochromatic Radiation

Einstein's derivation:

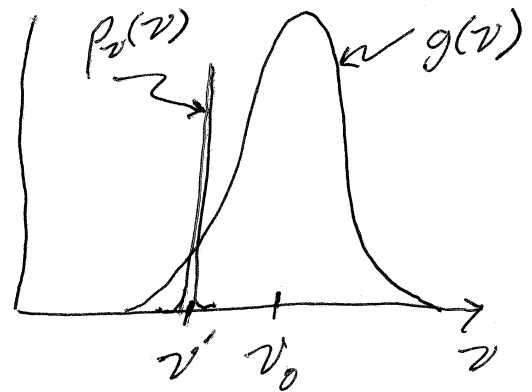
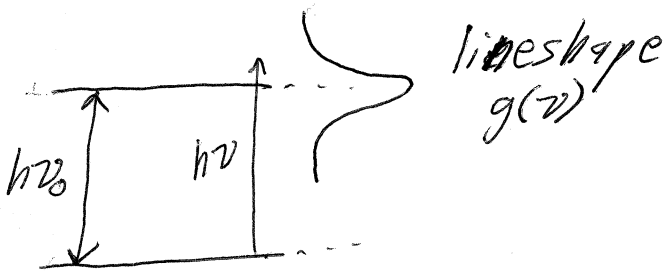


$$W_{12}^{\text{ind}} = B_{12} \rho_{\nu}(\nu_0)$$



Must generalize

$$W_{12} = \int B_{12} \rho_{\nu}(\nu) g(\nu) d\nu$$



For $\rho_{\nu}(\nu)$ much narrower than $g(\nu)$,

$$W_{12} \approx B_{12} g(\nu') \int \rho_{\nu}(\nu) d\nu$$

$$\boxed{W_{12} \approx B_{12} g(\nu') \rho}$$

$$\rho \equiv \int \rho_{\nu}(\nu) d\nu$$

Now $\rho = \frac{\text{total energy}}{\text{Vol}}$

before $\rho_{\nu} = \frac{\text{energy}}{\Delta\nu \cdot \text{Vol}}$

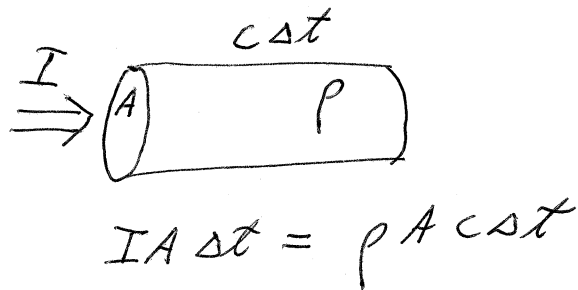
Other limit reduces to Einstein's case:

$$W_{12} \approx B_{12} \rho(\nu_0) \underbrace{\int g(\nu) d\nu}_{\approx 1 \text{ normalized}}$$

Intensity

$$I = c \rho$$

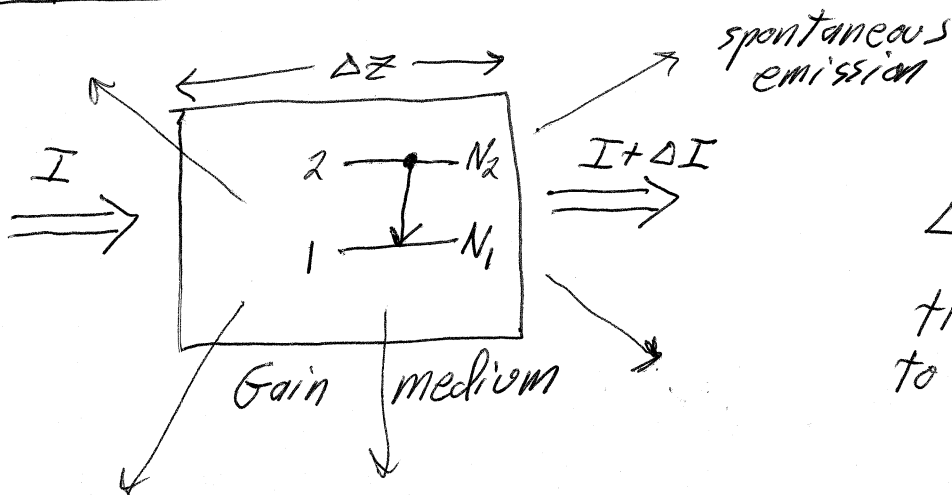
\uparrow energy / area · time \uparrow energy / Vol.



So

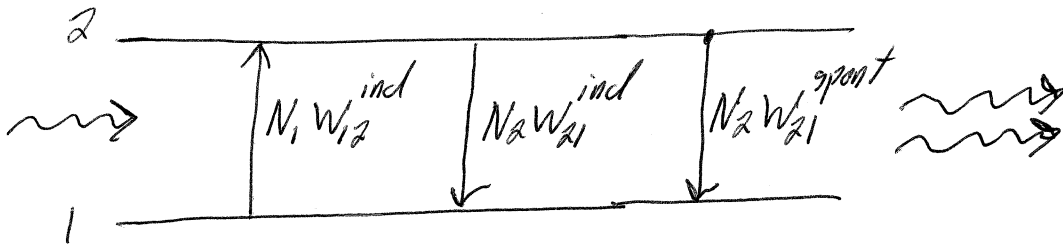
$$W_{12} = B_{12} g(\nu') \frac{I}{c}$$

Amplification by Atomic System



$$\Delta t = \frac{\Delta z}{c}$$

time for light
to pass through Δz



$$\frac{\Delta \rho}{\Delta t} = \frac{\text{energy generated}}{(\text{unit time})(\text{Vol})} = h\nu (N_2 W_{21}^{\text{ind}} - N_1 W_{12}^{\text{ind}})$$

But since $B_{12} = B_{21}$, then $W_{12}^{\text{ind}} = W_{21}^{\text{ind}}$

so

$$\frac{\Delta \rho}{\Delta t} = h\nu W_{21}^{\text{ind}} (N_2 - N_1)$$

$$\frac{\Delta \rho}{\Delta t} = h\nu W_{21}^{\text{ind}} \Delta N$$

$\Delta N \equiv N_2 - N_1$
"population inversion"

$$\Delta I = c \Delta \rho = h\nu W_{21}^{\text{ind}} \Delta N \Delta z$$

Now use $W_{21}^{\text{ind}} = B_{21} g(\nu') \frac{I}{c}$

$$= A_{21} \frac{c}{8\pi h\nu} g(\nu') \frac{I}{c}$$

$$= A_{21} \frac{\lambda^2}{8\pi} g(\nu') I \frac{1}{h\nu}$$

To obtain

$$\Delta I = A_{21} \frac{\lambda^2}{8\pi} g(\nu') I \Delta N \Delta z$$

Take limit as $\Delta z \rightarrow 0$

$$\boxed{\frac{dI}{dz} = \gamma I}$$

$$\boxed{\gamma(\nu) = A_{21} \frac{\lambda^2}{8\pi} g(\nu) \Delta N}$$

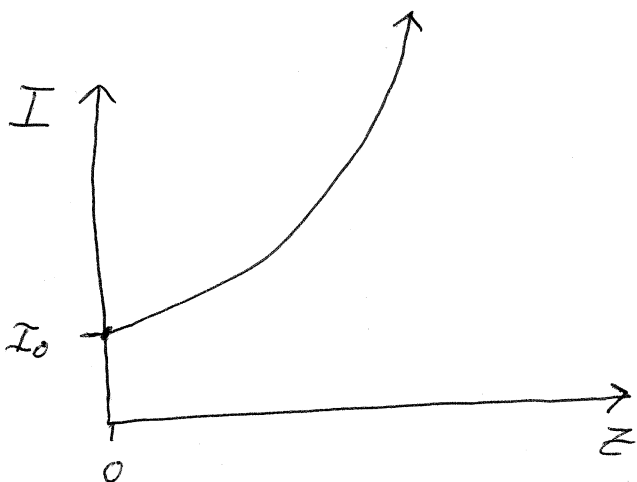
gain coefficient

In gain medium with refractive index n ,
 $\lambda \rightarrow \lambda/n$ so

$$\delta(\nu) = A_{21} \frac{\lambda^2}{8\pi n^2} g(\nu) \Delta N$$

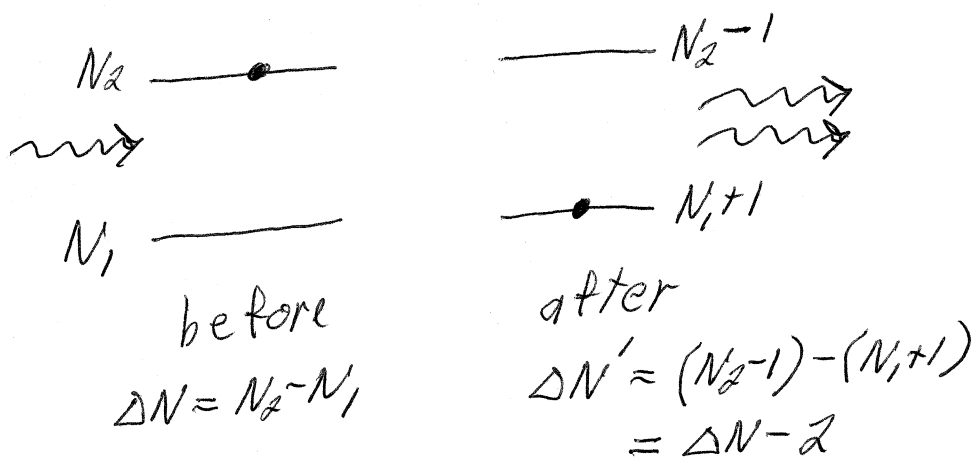
If $\delta = \text{constant}$, solution for $I(z)$ is

$$I(z) = I(0) e^{\delta z}$$



goes to ∞ : not realistic

Actually, δ not constant because ΔN not constant



This is called "saturation" of the gain

Absorption and Emission Cross Section

$$\delta(\nu) \equiv \Delta N \sigma(\nu) = (N_2 - N_1) \sigma(\nu)$$

gain coefficient = (population difference)(cross section)

$$\sigma(\nu) = A_{21} \frac{\lambda^2}{8\pi\nu^2} g(\nu)$$

depends on properties of a single atom

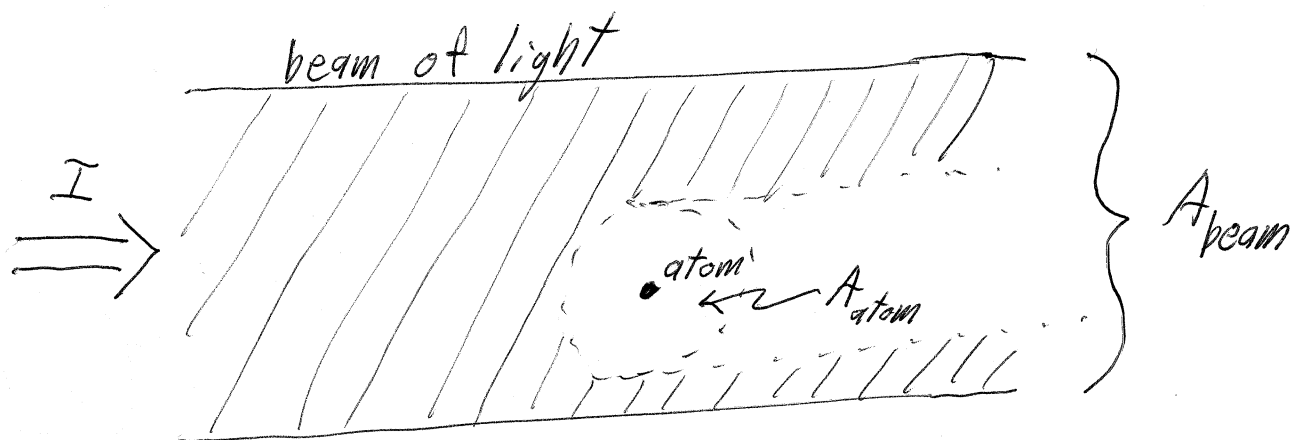
This cross section can be interpreted as the effective area within which an atom absorbs or emits light.

Consider $N_2 = 0$ (no excited atoms)

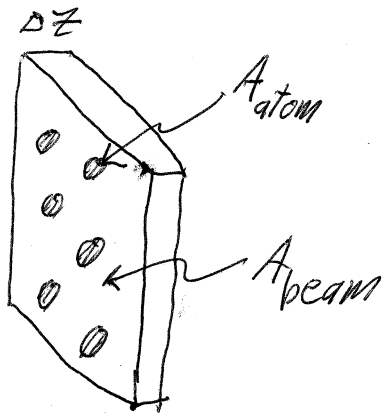
$$\delta = -N_1 \sigma$$

Define $\alpha \equiv N_1 \sigma$ absorption coefficient

Then $I(z) = I(0) e^{-\alpha z}$ (Beer's law)



Each atom absorbs light within an area A_{atom}
 Fraction of light absorbed by each atom is $\left(\frac{A_{atom}}{A_{beam}}\right)$



$$\begin{aligned} \# \text{ atoms in } \Delta z \text{ is} \\ = N_1 A_{\text{beam}} \Delta z \end{aligned}$$

Fraction of light absorbed in Δz is

$$\text{frac} = (N_1 A_{\text{beam}} \Delta z) \left(\frac{A_{\text{atom}}}{A_{\text{beam}}} \right) = N_1 A_{\text{atom}} \Delta z$$

But from definition of gain coefficient,

$$\frac{\Delta I}{I} = -\alpha \Delta z \quad \text{so} \quad \text{frac} = \alpha \Delta z$$

Therefore,

$$\alpha \Delta z = N_1 A_{\text{atom}} \Delta z$$

$$N_1 \sigma = N_1 A_{\text{atom}}$$

$$\boxed{\sigma = A_{\text{atom}}}$$

Cross section $\sigma(\nu)$:

- * property of single atom
- * varies with ν or λ [lineshape function $g(\nu)$]
- * can look up $\sigma(\nu)$ data in handbooks
- * same shape as lineshape function $g(\nu)$