WORCESTER POLYTECHNIC INSTITUTE MECHANICAL ENGINEERING DEPARTMENT

Optical Metrology and NDT ME-593L/ME-5304, C'2025

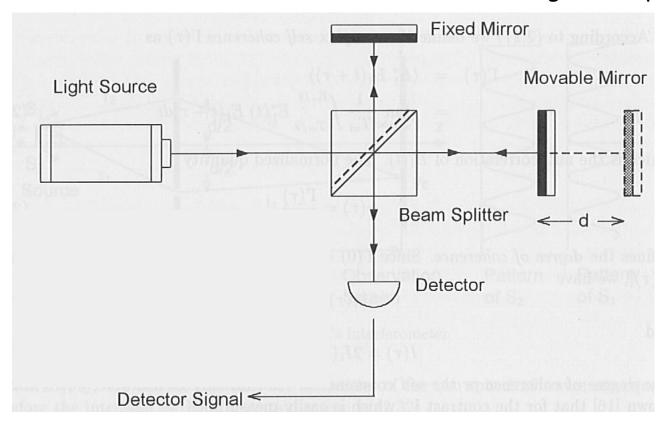
Introduction: Wave Optics February 2025





Temporal coherence

- Review interference equation: Lecture 04
- Michelson interferometer to test for coherence length (temporal)



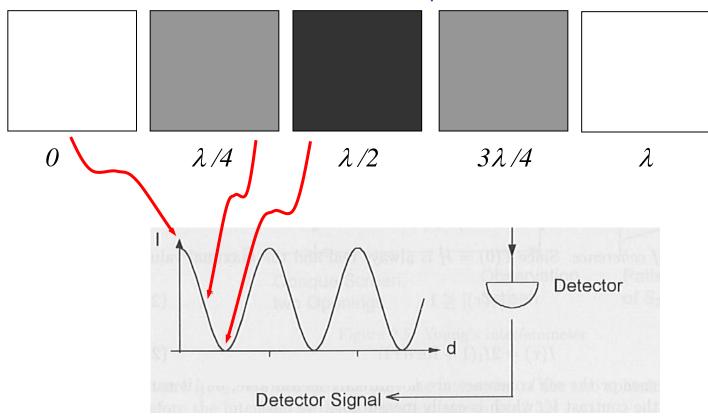




Temporal coherence

Assume 100% flat and orthogonal mirrors

Detected intensities as a function of position of movable mirror







Temporal coherence

Temporal coherence describes the self-correlation of a wave.
 Consider the normalized autocorrelation function

$$g(\tau) = \frac{G(\tau)}{G(0)} = \frac{\left\langle U^*(t) U(t+\tau) \right\rangle}{\left\langle U^*(t) U(t) \right\rangle}, \tag{1}$$

where τ is the time delay, related optical path length changes introduced by the movable mirror.

Equation 1 is called the degree of temporal coherence with

$$0 \le |g(\tau)| \le 1. \tag{2}$$

Usually, $|g(\tau)|$ drops from its largest values $|g(\tau)|=1$ as τ increases and the fluctuations become uncorrelated for sufficiently large time delay τ .





Temporal coherence

Coherence length is defined as

$$l_c = c \, \tau_c \;, \tag{3}$$

where c is the speed of light and τ_c is the power equivalent width of the function $|g(\tau)|$ given as

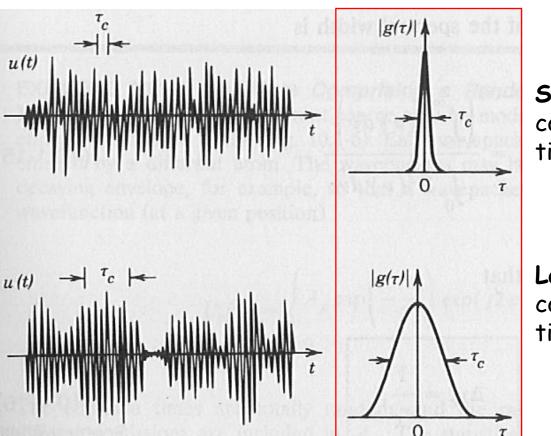
$$\tau_c = \int_{-\infty}^{\infty} |g(\tau)|^2 d\tau \,. \tag{4}$$





Temporal coherence

Magnitude of the degree of temporal coherence $|g(\tau)|$



Short coherence time

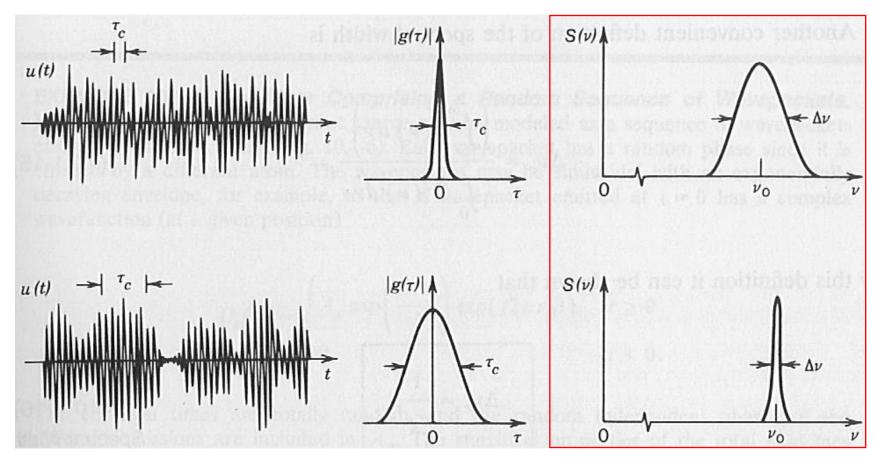
Long coherence time





Temporal coherence

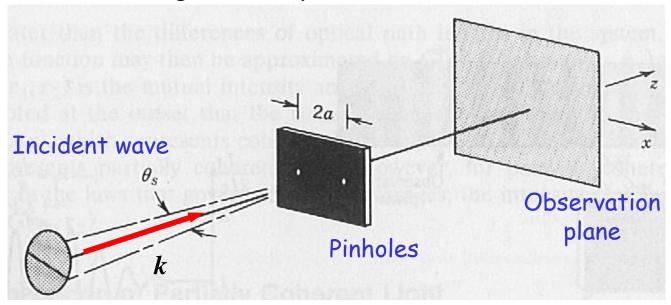
Power spectral density:
$$S(v) = \int_{-\infty}^{\infty} g(\tau) \exp(-j2\pi v \tau) d\tau$$



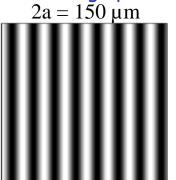


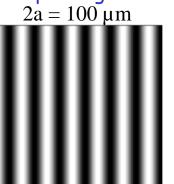


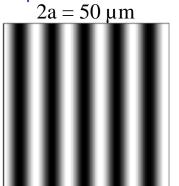
Spatial coherence Young's double pinhole interferometer



Predicted fringe patterns corresponding to difference pinholes distances



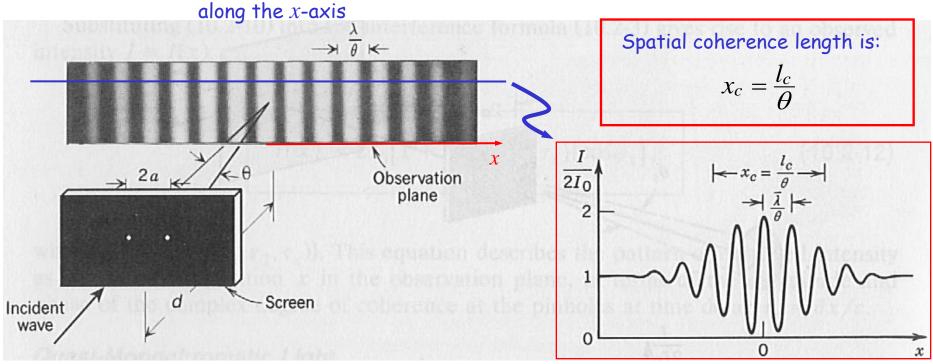




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Spatial coherence

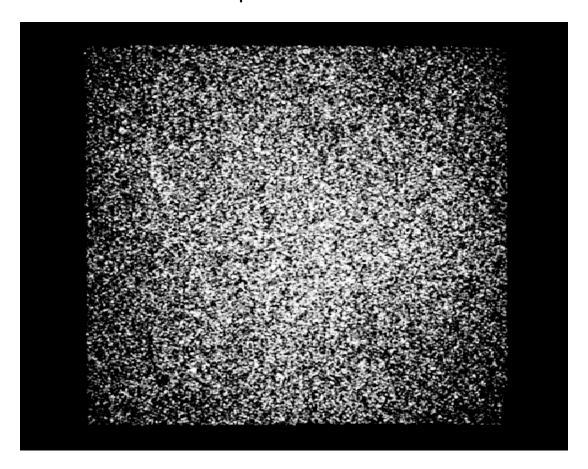
Actual fringe pattern: note variation of intensity







Speckle pattern observed on the flat surface of an object illuminated with a coherent light source with characteristic wavelength λ = 994.1 nm and horizontally polarized. Average surface roughness is 5 μ m, object distance is 30 mm, and f/2.8 aperture is used. The surface is illuminated and observed normal to the plane where the surface lies.

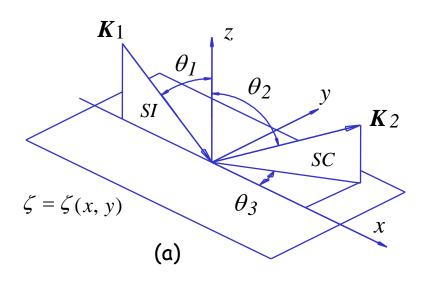


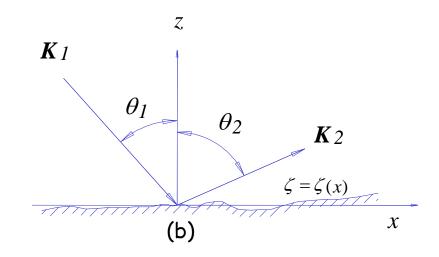




Illumination and scattering geometry from surfaces defined as $\zeta = \zeta(x, y)$ and $\zeta = \zeta(x)$.

Vectors of illumination and scattering: $\mathbf{K}_1 = k \ \hat{K}_1$, $\mathbf{K}_2 = k \ \hat{K}_2$









Scattering from a smooth surface

 $lue{}$ The complex amplitude of the scattered light, F_o , at point p, can be predicted using the Kirchoff integral theorem

$$F_o = \frac{1}{4\pi} \left\{ \iint_{S} \frac{1}{r} \exp(-jkr) \nabla U \cdot d\mathbf{S} - \iint_{S} U \nabla \left[\frac{1}{r} \exp(-jkr) \right] \cdot d\mathbf{S} \right\}$$
 (5)

which is derived from the Helmholtz equation.

□ For a smooth surface, $\zeta(x) \approx 0$, extending from [-L, L]:

$$F_o = F_o(\theta_1, \theta_2) = \frac{\sin[kL(\sin\theta_1 - \sin\theta_2)]}{kL(\sin\theta_1 - \sin\theta_2)}$$

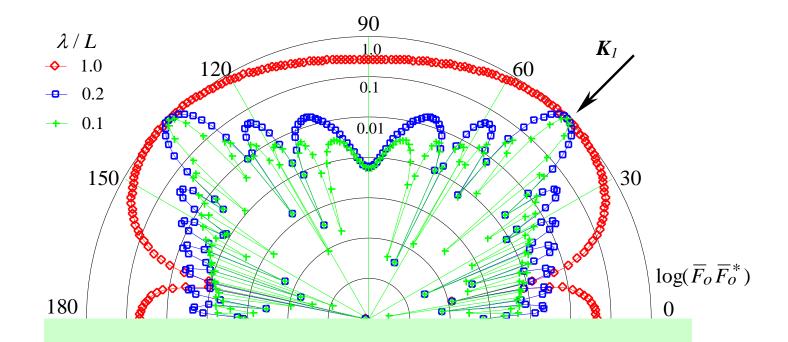
$$= \sin c \left[\frac{2\pi}{\lambda} L(\sin\theta_1 - \sin\theta_2) \right]$$
(6)





Scattering from a smooth surface

- $lue{}$ Scattering diagram for a smooth surface characterized by the ratio λ/L .
- $\hfill\square$ Note that when $\lambda/L \to 0$ scattering of light concentrates in the direction of specular reflection



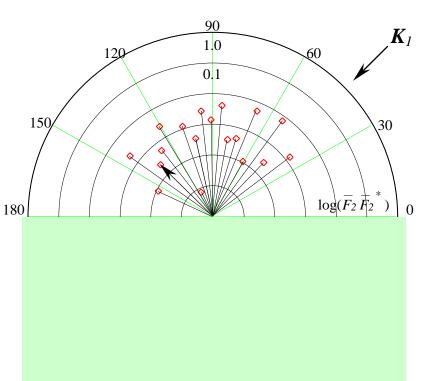


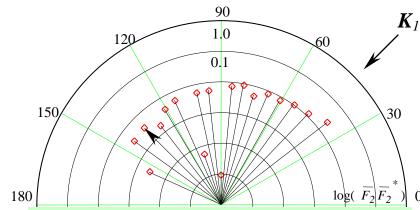


Scattering from a surface characterized by a periodic function:

$$\Lambda = 10\lambda$$
, $\theta_l = \pi/4$, $\lambda/h = 0.033$

$$\Lambda = 10\lambda$$
, $\theta_1 = \pi/4$, $\lambda/h = 0.066$





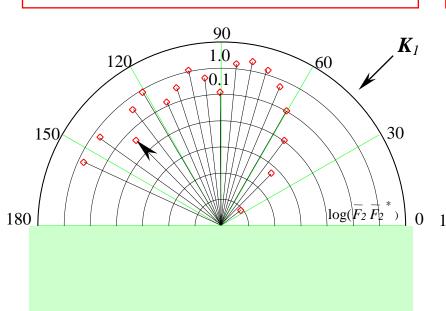


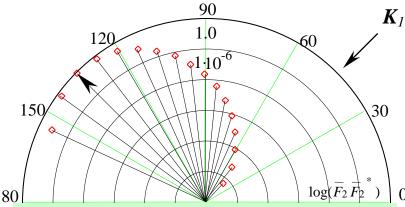


Scattering from a surface characterized by a periodic function

$$\Lambda = 10\lambda$$
, $\theta_I = \pi/4$, $\lambda/h = 1.0$

$$\Lambda = 10\lambda$$
, $\theta_I = \pi/4$, $\lambda/h = 10.0$



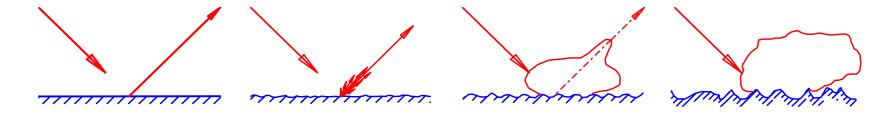






Scattering from a surface characterized by a random function

Transition from specular to diffuse scattering reflection as a function of surface roughness:

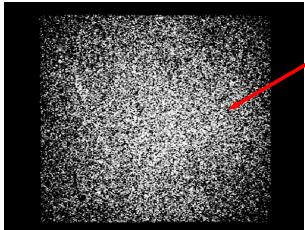




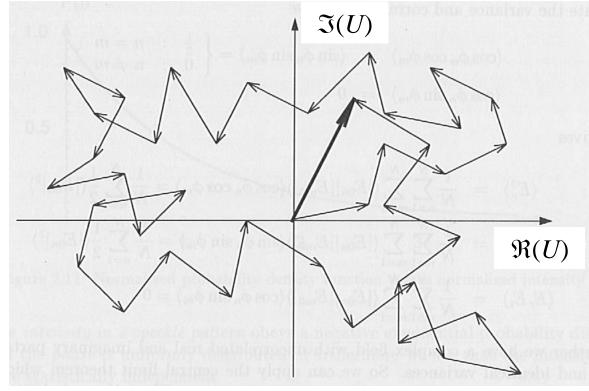


(x,y,z)

Speckle pattern



Adding contribution of scattered components at point (x,y,z)







Speckle properties: first order statistics

 $lue{}$ Sum of N complex amplitude components:

$$U(x, y, z) = \frac{1}{\sqrt{N}} \sum_{k=1}^{N} |A_k| \exp[j\phi_k(x, y, z)]$$
 (7)

Real an imaginary components:

$$\Re\{U\} = \frac{1}{\sqrt{N}} \sum_{k=1}^{N} \left| a_k \right| \cos(\phi_k)$$

$$\Im\{U\} = \frac{1}{\sqrt{N}} \sum_{k=1}^{N} \left| a_k \right| \sin(\phi_k)$$
(8)

Amplitude and phase are statistically independent.





Speckle properties: first order statistics

□ Probability distribution for the intensity is (mean = $2\sigma^2$, variance = $\langle I^2 \rangle$)

$$P_{I}(I) = \int_{-\pi}^{\pi} P_{I,\phi}(I,\phi)d\phi = \begin{cases} \frac{1}{2\sigma^{2}} \exp\left(-\frac{I}{2\sigma^{2}}\right) & ; I \ge 0\\ 0 & ; \text{ otherwise} \end{cases}$$
(9)

Probability distribution for the phase is:

$$P_{\phi}(\phi) = \int_{0}^{\infty} P_{I,\phi}(I,\phi)dI = \begin{cases} \frac{1}{2\pi} & ; -\pi \le \phi \le \pi \\ 0 & ; \text{ otherwise} \end{cases}$$
 (10)



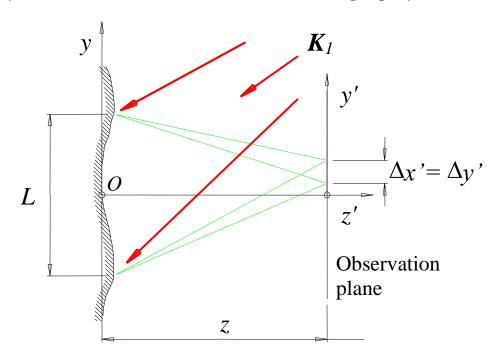


Speckle properties: second order statistics

Observed speckle size is (without imaging system):

$$\delta x_S = 2 \frac{\lambda z \pi}{L} \tag{11}$$

Speckle field formation without imaging system







Speckle properties: second order statistics

Observed speckle size is (with imaging system):

$$\delta r_{S} = 2.44 \frac{\lambda z}{D} \tag{12}$$

Speckle field formation with imaging system

