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http://www.wpi.edu/~ck/teaching1.htm
 → PH1130, C08

I: Length Contraction

$$l = l_0 \sqrt{1 - \frac{u^2}{c^2}}$$

$$l' = l \sqrt{1 - \frac{u^2}{c^2}}$$

prime = denotes physical quantities in moving reference frame

unprimed = stationary reference frame

$$\gamma = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}}$$

c = speed of light = 3.00×10^8 m/s
 u = velocity of moving frame (example: car)

II: Time Dilation (page 1275)

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{u^2}{c^2}}}$$

$$\Delta t' = \sqrt{1 - \frac{u^2}{c^2}} \Delta t = \gamma(\Delta t)$$

37.61

SUV

moon

spaceship: 0.8c past your SUV

a.) You measure spaceship travels 1.2×10^8 m
 t on race pilot = ? (find $\Delta t'$)

cont. →
 back

$$\Delta t = \frac{\Delta t'}{\gamma} = \frac{d}{u\gamma} = \frac{1.20 \times 10^8}{0.8c\gamma}$$

$$\Delta t_0 = 0.300 \text{ s}$$

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Possibly Helpful Formulas:

$$E^2 = \underbrace{(mc^2)^2}_{\text{rest energy}} + \underbrace{(pc)^2}_{\text{kinetic}}$$

total energy

zero rest mass. $E = pc$

$$\vec{p} = \frac{m\vec{v}}{\sqrt{1 - \frac{v^2}{c^2}}} = \gamma m\vec{v}$$

$$K = (\gamma - 1) mc^2$$

$$K = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}} - mc^2$$

$$\Delta t = \frac{\Delta t'}{\sqrt{1 - \frac{u^2}{c^2}}} = \gamma \Delta t' \quad (\text{p. 1298})$$

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Possibly Helpful Formulas:

$$\Delta t' - \Delta t = \frac{\Delta t (g) \Delta h}{c^2}$$

$$E = \frac{mc^2}{\sqrt{1 - v^2/c^2}}$$

$$f = \sqrt{\frac{1 - v/c}{1 + v/c}} \text{ for moving}$$

$$g = 10 \text{ m/s}^2$$

$$T = 1/f$$

$$\sum \vec{F} = m\vec{a}$$

$$\Delta t' - \Delta t = \left(-\frac{1}{2} \left(\frac{v}{c} \right)^2 + \frac{g \Delta h}{c^2} \right) \Delta t$$

* time is shorter when it is closer to the gravitational field

$$1 + x + x^2 + x^3 + \dots = \frac{1}{1-x} \quad |x| \ll 1$$

$$1 + 2x + 3x^2 + \dots + nx^{n-1} = \frac{1}{(1-x)^2} \quad x \rightarrow -x \Rightarrow 1 - x + x^2 - x^3 + \dots = \frac{1}{1+x}$$

$$(1-x)^{-n} \approx 1 + nx$$

$$\text{nano} = 10^{-9}$$

Conference Outline:

1.) Photoelectric Effect:

$$K_{\max} = \begin{cases} 0 & f \leq f_0 \\ h(f - f_0) & f \geq f_0 \end{cases}$$

$$E = hf \quad h = 6.626 \times 10^{-34} \text{ Js}$$

Work function = Energy binded electrons needed to overcome = hf_0

2.) Conservation of Energy: $hf = hf_0 + K$

$$E^2 = (mca)^2 + (K)^2 \quad \boxed{E = \gamma mca}$$

(K = $(\gamma - 1)mca$)

Work function Kinetic Energy

3.) Classic Mechanics: $\sum F = ma$

$$F_e = F_c \Rightarrow \frac{ke^2}{r} = \frac{m_e v^2}{r} \quad (k = \frac{1}{4\pi\epsilon_0} = 9 \times 10^9 \frac{\text{Nm}^2}{\text{C}^2})$$

$$\Rightarrow v = \sqrt{\frac{ke^2}{m_e r}}$$

$$E = \frac{1}{2} m_e v^2 - \frac{ke^2}{r} = -\frac{ke^2}{2r}$$

4.) Bohr Quantization: $m_e v r = n\hbar$ or $L = n \frac{h}{2\pi}$

$$L = \text{angular momentum} = m_e v r \quad \hbar = \frac{h}{2\pi}$$

$$r_n = \frac{n^2 \hbar^2}{ke^2 m_e} \text{ (radii)} \quad r_1 = 0.529 \text{ \AA} \quad (1 \text{ \AA} = 10^{-10} \text{ m})$$

$$E_n = -\frac{13.6 \text{ eV}}{n^2} \quad 1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$$

$$\text{conservation of Energy: } \Delta E = hf \Rightarrow hf = 13.6 \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

DeBroglie particle/wave duality:

$$\lambda = \frac{h}{p} = \frac{h}{\gamma m v} = \frac{h}{m v \sqrt{1 - \frac{v^2}{c^2}}}$$

$$f = \frac{E}{h} = \frac{\gamma m c^2}{h} = \frac{m c^2}{h \sqrt{1 - \frac{v^2}{c^2}}}$$

$$\boxed{E = pc} \\ \text{photon} \\ \text{Energy}$$

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Conference Formulas:

$$\lambda = \frac{h}{p} \quad h = 6.626 \times 10^{-34} \text{ Js}$$

perfect constructive interference:
 $\Delta x = n\lambda$
 \rightarrow difference in distance

accelerating electrons: $K = e \Delta V$
 non-relativistic electrons: $\lambda = \frac{h}{mv}$ $v = \sqrt{\frac{2K}{m}}$

Relativistic electrons: $\lambda = \frac{h}{\gamma mv}$ $E = K + mc^2$
 $p = \sqrt{\frac{E^2}{c^2} - (mc)^2}$
 $\lambda = \frac{h}{p} = \frac{h}{\sqrt{\frac{K^2}{c^2} + 2Kmc}}$

E-field of wave: $\Psi \sim a \Psi_0 \cos\left(\frac{\pi d \sin \theta}{\lambda}\right)$
 Intensity = $4I_0 \cos^2\left(\frac{\pi d \sin \theta}{\lambda}\right)$

$|\Psi|^2 \rightarrow$ intensity of wave gives probability that an individual electron would strike the screen at a given place

$$P(x,t) = |\Psi(x,t)|^2 dx$$

Running wave $\rightarrow \Psi(x,t) = \Psi_0 e^{i(kx - \omega t)}$
 $\int_{-\infty}^{\infty} P(x,t) = 1 \Rightarrow \int_{-\infty}^{\infty} |\Psi(x,t)|^2 dx = 1$

Compton Effect:

Compton scattering equation:

$$\lambda' - \lambda = \frac{h}{m_e c} (1 - \cos \theta)$$

λ = wavelength of photon before scattering

λ' = wavelength of photon after scattering

$$\frac{1}{f'} - \frac{1}{f} = \frac{h}{m_e c^2} (1 - \cos \theta)$$

hydrogen atom

$$\Delta E = 13.6 \text{ eV} \left(\frac{1}{n_i^2} - \frac{1}{n_f^2} \right)$$

$$m_e = 9.11 \times 10^{-31} \text{ kg}$$